2nd COPIGAL Workshop on Studies of Exotic Nuclei J. Skalski (NCBJ)

M. Bender (CENBG)

Correspondence between self-consistent and micro-macro predictions for very deformed & spin isomeric configurations.

Coworkers: M. Kowal, W. Brodzinski (NCBJ), P. Jachimowicz (UZ); Benoit Avez, Benjamin Bally, Jeremy Sadoudi (CENBG) Exotic shapes and configurations in heavy & superheavy nuclei (SHN)

- These configurations may have larger halflives and show particular patterns allowing for their experimental verification.
- In particular, spin isomers in SHN may be longer-lived.
- Orbits seen at super- and hyperdeformation in actinides are those occupied at normal shape in SHN; they can provide a test of a model. Experiments confirming predicted minima validate nuclear models.

# Good methods should give similar predictions.

- Micro-macro, as a simpler one, is better tested/fitted against various data, eg. fission half-lives.
- Selfconsistent methods could (if constructed properly) give better extrapolations.But it is not guarateed at present. Hence, a prudent idea is to see whether both methods give similar results.

IIIrd minima in actinides do they exist?

M. Kowal, J. Skalski PRC, in press.

Planned activity:

- more detailed determination of 3rd barriers around Th (including odd nuclei);
- study of oblate configurations in SHN, including decay hindrance from Kisomerism;
- Importance of beyond-mean-field effects.

Polikanov et al. (1962) Discovery of isomeric fission
Strutinsky (1967) Calculated second minima
Specht et al. (1972) Identification of the rotational band with large moment of inertia in 240Pu



Third minima: Th,U First predicted: P. Moller, S.G. Nilsson and R.K. Sheline (1972) then Howard & Moller (1980) – rather shallow III-rd minima S.Cwiok et al. – rather deep III-rd minima some, not all, HF calculations give III-rd minima, BUT

they often differ from macro-micro results

Experiments:

1)Studies of microstructure in the resonances of fission probability found using (n,f), (t,pf) and (d,pf) reactions
B.B. Back et al. (1972)
J. Blons et al. (1975)
recent claims of III-rd minima in 232,234,236U
2) Also observations of asymmetric angular distribution of light fission fragments around 232Th

## TKE



TKE as a function of the mass distribution of the fission fragments.

M. Csatlos et. al., Acta Phys. Pol. B Vol. 34, (2003).



P.G. Thirolf, D. Habs / Prog. Part. Nucl. Phys. 49 (2002) 325-402

#### Macroscopic-microscopic approach:

$$E = E_{tot}(\beta_{\lambda\mu}) - E_{MACRO}(\beta_{\lambda\mu}=0)$$

## $E_{tot}(\beta_{\lambda\mu}) = E_{MACRO}(\beta_{\lambda\mu}) + E_{MICRO}(\beta_{\lambda\mu})$

### $\circ$ E<sub>MACRO</sub>( $β_{\lambda\mu}$ ) = Yukawa + exp

### $\odot E_{MICRO}(\beta_{\lambda \mu}) = Woods - Saxon + pairing BCS$

## Shape Parametrization:

 $R(\Theta, \Phi) = \left\{ 1 + a_{20}Y_{20} + a_{40}Y_{40} + a_{60}Y_{60} + a_{80}Y_{80} \right\}$  $+a_{22}Y_{22}^{(+)}+a_{42}Y_{42}^{(+)}+a_{AA}Y_{AA}^{(+)}$  $+a_{22}Y_{22}^{(+)}+a_{52}Y_{52}^{(+)}$  $+a_{30}Y_{30}+a_{50}Y_{50}+a_{70}Y_{70}$  $Y_{\lambda\mu}^{(+)} = \frac{1}{\sqrt{2}} (Y_{\lambda\mu} + Y_{\lambda-\mu})$ 





S.Cwiok at al



#### BÜRVENICH, BENDER, MARUHN, AND REINHARD PHYSICAL REVIEW C 69, 014307 (2004)



Careful calculations for 234U by M. Bender do not show any 3rd minimum.

## Performance of our model (HN)

- □ Second barriers in actinides,
- □ Second minima in actinides,
- □ Fission barriers for SHN



Z	Ν	Α	LSD	FRLDM	HN	EXP1	EXP2	Theoretical models: LSD FRLDM HN
90	136	226	_	7.20	6.26	_	_	Experimental data: [23] [24] [23] [24] [23] [24]
	138	228	_	6.53	5.92	_	6.5	N 19 18 14 99 14 99
	140	230	_	5.65	5.97	6.80	6.1	
	142	232	5.2	5.45	6.07	6.70	6.2	$< B_{f}^{in} - B_{f}^{exp}  > 0.78 \ 0.84  \ 0.79 \ 0.90  \ 0.39 \ 0.33$
	144	234	4.8	5.37	6.07	_	6.3	$Max \mid B_{\ell}^{th} - B_{\ell}^{exp} \mid 1.50 \mid 1.50 \mid 1.85 \mid 2.33 \mid 0.83 \mid 0.86$
	146	236	-	6.04	6.35	_	_	5 000 004 005 111 046 040
92	138	230	_	4.28	5.64	_	_	$\sigma_{\rm RMS} = 0.92 \ 0.94 \ 0.95 \ 1.11 \ 0.46 \ 0.40$
	140	232	_	4.73	5.66	5.40	5.3	
	142	234	4.7	4.89	5.84	5.50	5.7	
	144	236	4.3	5.03	5.78	5.67	5.6	
	146	238	4.9	5.64	5.93	5.50	5.8	$\frac{248}{5}$ 0.0 $\frac{248}{5}$ $\frac{248}{5}$ $\frac{248}{5}$ $\frac{248}{5}$ $\frac{248}{5}$ $\frac{248}{5}$ $\frac{248}{5}$ $\frac{248}{5}$
	148	240	4.6	6.37	6.04	_	5.8	$\dot{\underline{U}} = 0.2$ $^{232}U$ $^{236}U$ $^{240}Pu$ $^{246}Cm$
)	150	242	_	7.10	6.23	_	_	Δ <sub>0.0</sub>
94	140	234	_	-	4.68	_	_	€ -0.2 <sup>24</sup> #Cm
	142	236	4.4	4.36	5.06	_	4.5	= -0.4
	144	238	5.3	4.47	5.15	5.10	5.2	-0.6 2 <sup>34</sup> Th
	146	240	4.5	4.91	5.28	5.15	5.3	-0.8 - <sup>232</sup> Th
	148	242	5.0	5.72	5.52	5.05	5.3	136 138 140 142 144 146 148 150 152 154 156
	150	244	6.1	6.47	5.63	4.85	5.2	N
	152	246	6.6	7.07	5.75	_	5.3	
	154	248	_	-	5.42	_	_	
96	144	240	_	3.92	4.25	_	_	
	146	242	4.2	4.45	4.42	5.00	4.0	1.0
	148	244	5.3	5.07	4.72	5.10	4.3	0.8
	150	246	5.2	5.87	4.98	4.80	4.7	α 0.6 242 Cm <sup>242</sup> Cm <sup>240</sup> Ll 246 Pu <sup>250</sup> Cm
	152	248	6.3	6.65	5.14	4.80	5.0	$\times 0.4$ $^{46}\text{Cm}$ $^{236}\text{Pu}_{236}$ $^{238}\text{U}_{242}$ $^{248}\text{Cm}$ $^{248}\text{Cm}$
	154	250	5.3	6.25	4.87	_	4.4	$\mathbf{G} = 0 \mathbf{Z} \mathbf{U} \mathbf{U} \mathbf{U} \mathbf{U} \mathbf{U} \mathbf{U} \mathbf{U} U$
	156	252	_	5.68	4.31	_	_	$f = -0.2$ $2^{30}$ Th $2^{32}$ Th $2^{38}$ Pu
98	150	248	_	5.18	4.14	_	_	$\mathbf{\Omega} = -0.4$ $228_{\pm 1}$
	152	250	4.6	5.92	4.57	_	3.8	-0.6
	154	252	_	5.83	4.36	_	3.5	-0.8
	156	254	_	5.27	3.95	_	_	
								N



Ν	А	$E_{II}^{min}(th)^*$	$E_{II}^{min}(exp)^*$
144	236	2.04	2.75
146	238	1.94	2.56
142	236	2.43	3.00
144	238	2.05	2.40
146	240	1.95	2.80
148	242	1.99	2.20
144	240	1.69	2.00
146	242	1.64	1.90
148	244	1.68	2.20(?)

NIKOLOV, SCHUNCK, NAZAREWICZ, BENDER, AND PEI

PHYSICAL REVIEW C 83, 034305 (2011)

Ba	arrie	ers ir	ר S⊢	IN		126 124 <b>B</b> (MeV)
Nucleus	SHF	FRLDM	ETFSI	HN	EXP	122 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1
<sup>284</sup> 112 <sub>172</sub>	6.06	7.41	2.2	4.29	5.5	
<sup>286</sup> 112 <sub>174</sub>	6.91	8.24	3.6	5.01	5.5	
<sup>288</sup> 114 <sub>174</sub>	8.12	9.18	6.1	5.53	6.7	
$^{290}114_{176}$	8.52	9.89	6.6	5.83	6.7	
$^{292}114_{178}$	_	9.98	7.2	6.34	6.7	
<sup>292</sup> 116 <sub>176</sub>	9.35	9.26	6.5	6.22	6.4	132 136 140 144 148 152 156 160 164 168 172 176 180 184 188 192
<sup>294</sup> 116 <sub>178</sub>	9.59	9.46	7.2	6.28	6.4	N
$^{296}116_{180}$	_	9.10	7.2	6.07	6.4	126
$^{294}118_{176}$	_	8.48	6.6	5.99	_	$124 \begin{bmatrix} \delta B_{f}^{\gamma_{2}+a_{42}+a_{44}} \text{ (MeV)} \\ 122 \end{bmatrix} = 2.0 $
$296118_{178}$	_	8.36	7.0	6.04	_	
$^{298}118_{180}$	_	8.05	7.4	5.72	_	
<sup>296</sup> 120 <sub>176</sub>	_	7.69	6.2	5.64	_	
<sup>298</sup> 120 <sub>178</sub>	_	7.33	6.6	5.50	_	N 110
$^{300}12O_{180}$	_	7.01	6.8	5.05	_	
$^{302}120_{182}$	_	6.07	7.2	4.66	_	
<sup>304</sup> 120 <sub>184</sub>	_	4.86	6.8	4.20	_	102 100 98 2.0
						132 136 140 144 148 152 156 160 164 168 172 176 180 184 188 192
						N

### Back to IIIrd minima in actinides



The dipole deformation 1 is omitted there, as corresponding to a shift of the origin of coordinates which leaves energy (always calculated in the center of mass frame) invariant. However, this is true only for weakly deformed shapes. For large elongations, b1 acquires a meaning of a real shape variable.

## IIIrd minima – type: A



One can nd continuous 8D paths start ing at the supposed IIIrd minimum and leading to scission, along which energy decreases gradually.

minima with larger octupole deformations (A) have quadrupole moments Q 170 b, disturbingly close to the scission region.

minima (A) are just intermediate congurations on the scission path, whose energy was calculated erroneously because of limitations of the admitted class of shapes.

#### M. Kowal, J. Skalski PRL 2012 (in press) <u>arXiv:1203.4449</u>

#### IIIrd minima – type: B

 $\begin{array}{ll} -0.35 < \beta_1 < 0.00 & -0.55 < \beta_2 < 1.50 & 0.00 < \beta_3 < 0.35 \\ -0.10 < \beta_4 < 0.35 & -0.20 < \beta_5 < 0.20 & -0.15 < \beta_6 < 0.15 \end{array}$ 



# IIIrd minima – type: B en. difference from beta1=0:



## IIIrd minima – type: B



the barrier vanishes in uranium and must be smaller than 330 keV in 232Th. The only other nonzero upper limit on the IIIrd barrier of 200 keV we nd in 230Th.

#### So, the Woods-Saxon model does not predict deep IIIrd minima.



We study the existence of third, hyperdeformed minima in a number of even-even Th, U and Pu nuclei using the Woods-Saxon microscopicmacroscopic model that very well reproduces first and second minima barriers in actinides. and fission Deep (3 4 MeV) minima found previously by Cwiok et al. are found spurious after sufficiently general shapes are included. Shallow third wells may exist in 230Th,232Th, with IIIrd barriers < 200 and 330 keV (respectively). Thus, a problem of discrepancy qualitative between microscopic-macroscopic selfconsistent predictions and is resolved. Now, an understanding of experimental results on the apparent third minima in uranium becomes an issue.